Classical Mechanics 1, Autumn 2025 CMI

Assignment 8
Due by 6pm, 14 Nov, 2025
Kepler problem

1. $\langle \mathbf{3} + \mathbf{4} + \mathbf{4} \rangle$ CM and relative variables. The center of mass and relative vectors in the problem of 2 point particles of masses m_1 , m_2 and positions $\mathbf{r}_1, \mathbf{r}_2$ are defined as $\mathbf{R} = (m_1\mathbf{r}_1 + m_2\mathbf{r}_2)/M$ and $\mathbf{r} = \mathbf{r}_2 - \mathbf{r}_1$ where $M = m_1 + m_2$. Let \mathbf{p}_1 and \mathbf{p}_2 be the momenta of the two particles. Let us define $\mu_{1,2} = m_{1,2}/M$ and the reduced mass $m = m_1m_2/(m_1+m_2)$. (a) Draw a figure with a suitable origin (not at the center of mass) indicating the vectors $\mathbf{R}, \mathbf{r}, \mathbf{r}_1, \mathbf{r}_2$. (b) Show that

$$\mathbf{r}_1 = \mathbf{R} - \mu_2 \mathbf{r}, \quad \mathbf{r}_2 = \mathbf{R} + \mu_1 \mathbf{r}. \tag{1}$$

Let the CM momentum be defined as $P = M\dot{R}$ and the relative momentum $p = m\dot{r}$. Show that

$$P = p_1 + p_2, p_1 = \mu_1 p_2 - \mu_2 p_1 \text{ and } p_1 = \mu_1 P - p, p_2 = \mu_2 P + p.$$
 (2)

- (c) Express the total energy of the two-body central force problem (with potential V(r)) as a sum of center of mass and relative energies.
- 2. $\langle \mathbf{3} + \mathbf{2} + \mathbf{5} \rangle$ Recall that the (relative) energy in the Kepler problem may be expressed as $E = \frac{1}{2}m\dot{r}^2 + V_{\rm eff}(r)$ where $V_{\rm eff} = \frac{l^2}{2mr^2} \frac{\alpha}{r}$. Here r is the magnitude of the relative vector, m the reduced mass, $\alpha = Gm_1m_2$ and l the z component of relative angular momentum. We showed that for fixed $l \neq 0$ and E < 0, the orbit is an ellipse $r = \rho/(1 + \epsilon \cos \phi)$ (up to a possible rotation) where ρ is the radius of the circular orbit for that l (or semilatus rectum) and ϵ the eccentricity. (a) Find ρ in terms of l, m, α and also find the minimum value of the effective potential. (b) Express the length of the semimajor axis a in terms of ρ, ϵ . (c) Express the relative energy of this elliptical orbit in terms of ρ, ϵ and also in terms of the length of the semimajor axis a.